



Testing spin parity in galaxies, groups and clusters

Galaxy winding and projected group / cluster motions across the sky

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1 Why study spin parity across scales?

Does the cosmos favour a handedness?

Compare clockwise and anticlockwise outward winding across the sky.



Clockwise (CW)



Anticlockwise (ACW)

Mirror symmetry predicts no ensemble preference. Image winding is a proxy for spin, not a direct rotation measurement.

Could early physics leave a trace?

Primordial chirality -> halo spins

Simulations show that an initial handed pattern can survive in later halo spins [1].

Rotating black-hole universe (BHU)

A speculative proposal links a rotating parent black hole to a preferred axis and possible spin-direction asymmetry [2].

Rotation or an axis alone does not prove parity violation. We do not test these specific models.

Standard LambdaCDM: no preferred hand

Tidal torques can align spins [3].

Alignment can preserve mirror symmetry.

Correlated spins are allowed.

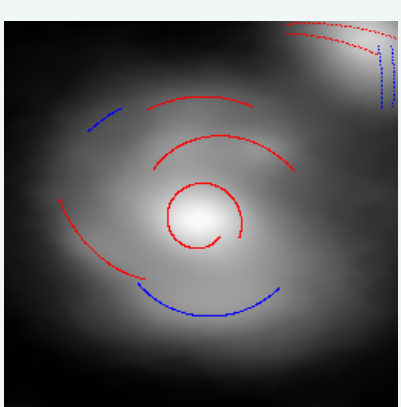
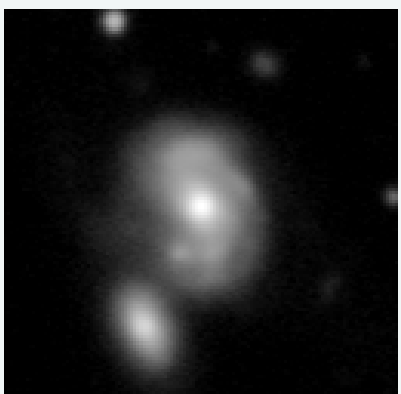
With parity-symmetric initial conditions, no ensemble handedness preference is expected. A finite sample can still be imbalanced.

Theory: [1] Shim et al. (2025 revision), arXiv:2406.06080; [2] Poplawski (2025 revision), arXiv:1910.10819; [3] Lee & Pen (2001), ApJ 555, 106. Full references: QR.

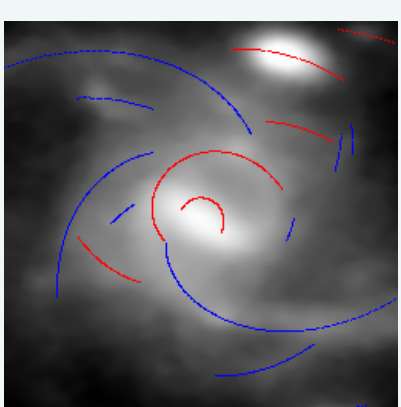
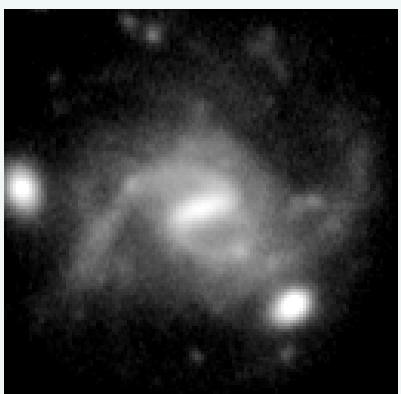
2 Galaxy winding: dipole as the main directional test

Our galaxies: images and fitted arms

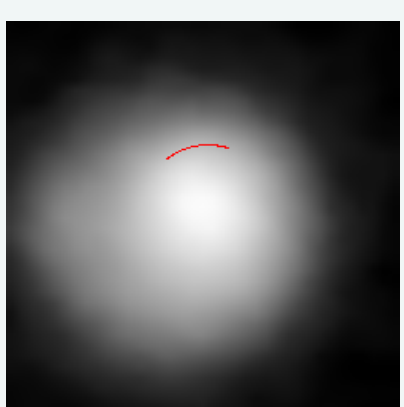
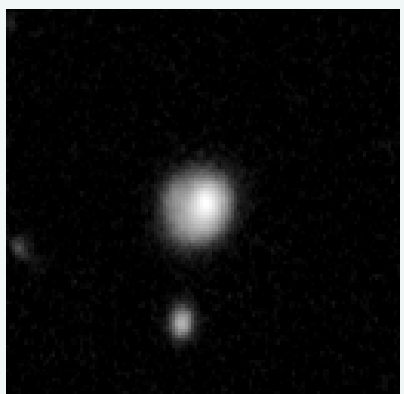
Brighter CW
r = 15.05



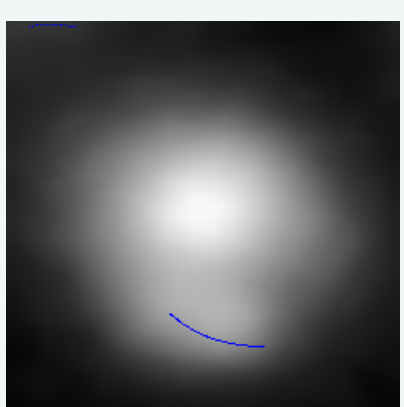
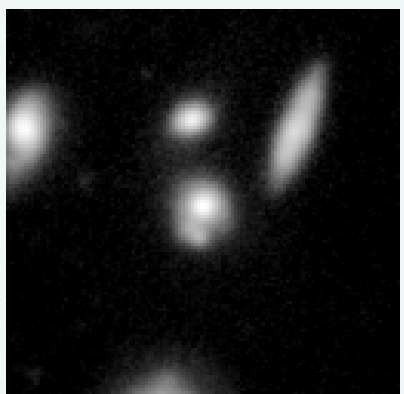
Brighter ACW
r = 15.96



Fainter CW
r = 17.64



Fainter ACW
r = 17.69



Our input cutouts (top) and stored arm fits (bottom).

Development examples, not validation. Brighter ACW chosen for visible structure. Curves are stored fits, not proof of spin. Smaller r = brighter.

30,910 resolved / 49,211 images

CW: 15,479

ACW: 15,431

Observed imbalance +0.155%

$A = (CW - ACW) / \text{resolved}; 18,301 \text{ unresolved}$

Selection: 46.7% of smallest vs 75.2% of largest resolve.

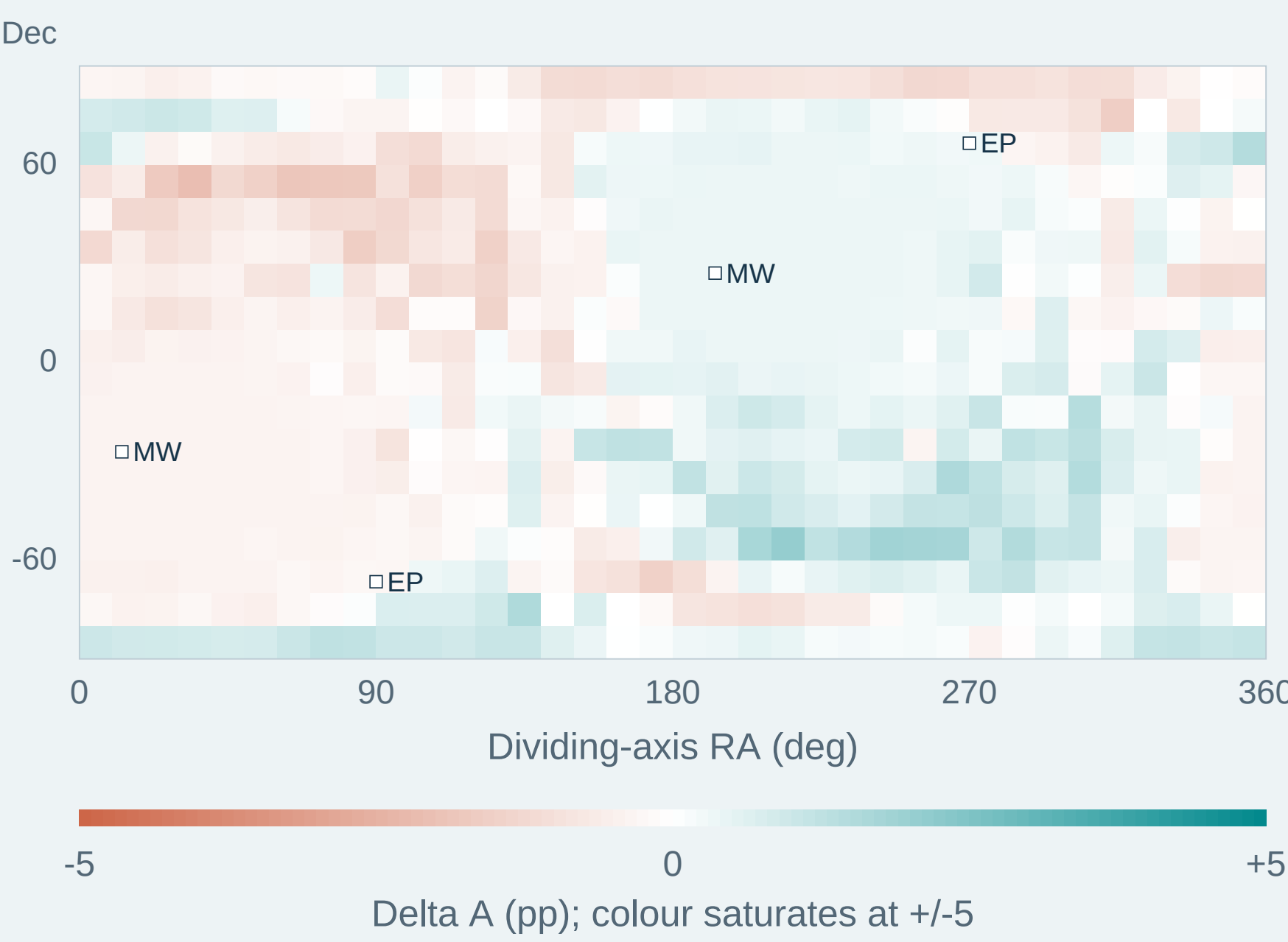
DESI: 7,905 positional matches; 4,920 also resolved.

DESI associations are provisional. Added dipole/quadrupole terms did not improve held-out prediction.

SpArcFiRe | 30,910 resolved

Stored spiral-arm fits; same 49,211 targets

Colour: imbalance toward the axis minus the opposite half.



CW 15,479 | ACW 15,431

Map cells: -2.1 to 2.1 pp (not a confidence range)

Historical dipole: 2.27%

Approx. 95% Wald: 0.00-7.79%

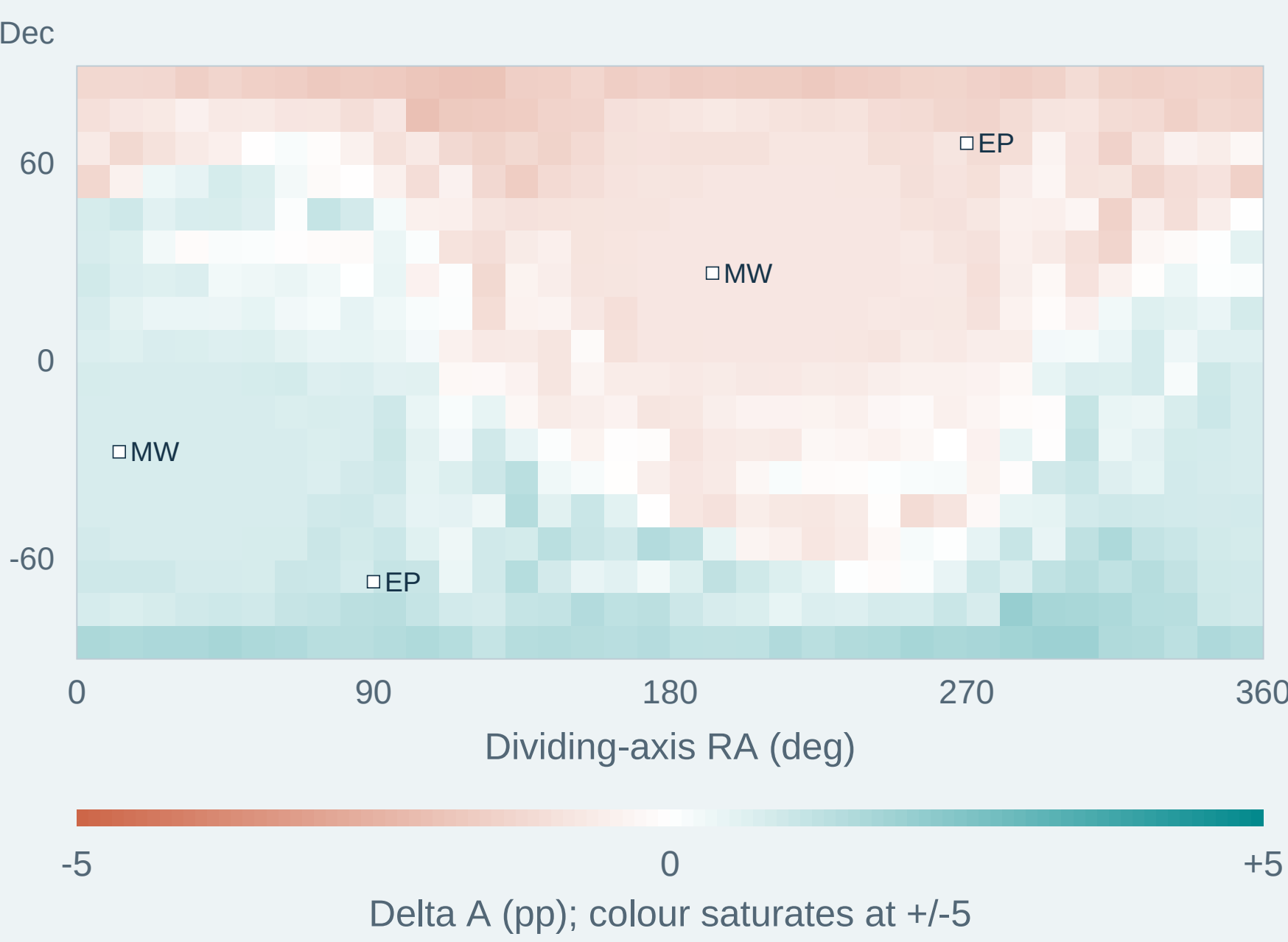
Conditional global shuffle $p = 0.684$

The map shows hemisphere contrast, not dipole amplitude. No preferred direction established.

Fourier | 49,007 assigned

Frozen image-pattern estimator; same targets

Colour: imbalance toward the axis minus the opposite half.



CW 24,644 | ACW 24,363

Map cells: -2.0 to 2.0 pp (not a confidence range)

Full dipole: 1.65% [0.00-6.03%]

Shared sample: 2.16% [0.00-7.68%]

Approx. 95% Wald envelopes; shared $n = 30,910$.

Four new tests: Bonferroni $p = 1.0$ each

Conditional shuffles; classification and selection biases remain. No population calibration transfer.

Two algorithms, one 49,211-target inventory; resolved sets differ. CW/ACW share one display orientation. Maps are observed contrasts, not detections. Both resolve 30,910 (18,791 agree). SpArcFiRe GZ/CzSL: -0.06% [-4.44%, +4.32%], assuming shared errors. Cohort differences: [-19.07%, +11.10%].

3 Groups and clusters: exploratory motion patterns

How we measure projected motion

1 Regression

Fit member line-of-sight velocity versus position angle using a sine and cosine.

Velocities -> angular fit -> direction

2 Half-plane

Rotate a dividing line; compare the two mean velocities and extract the first harmonic.

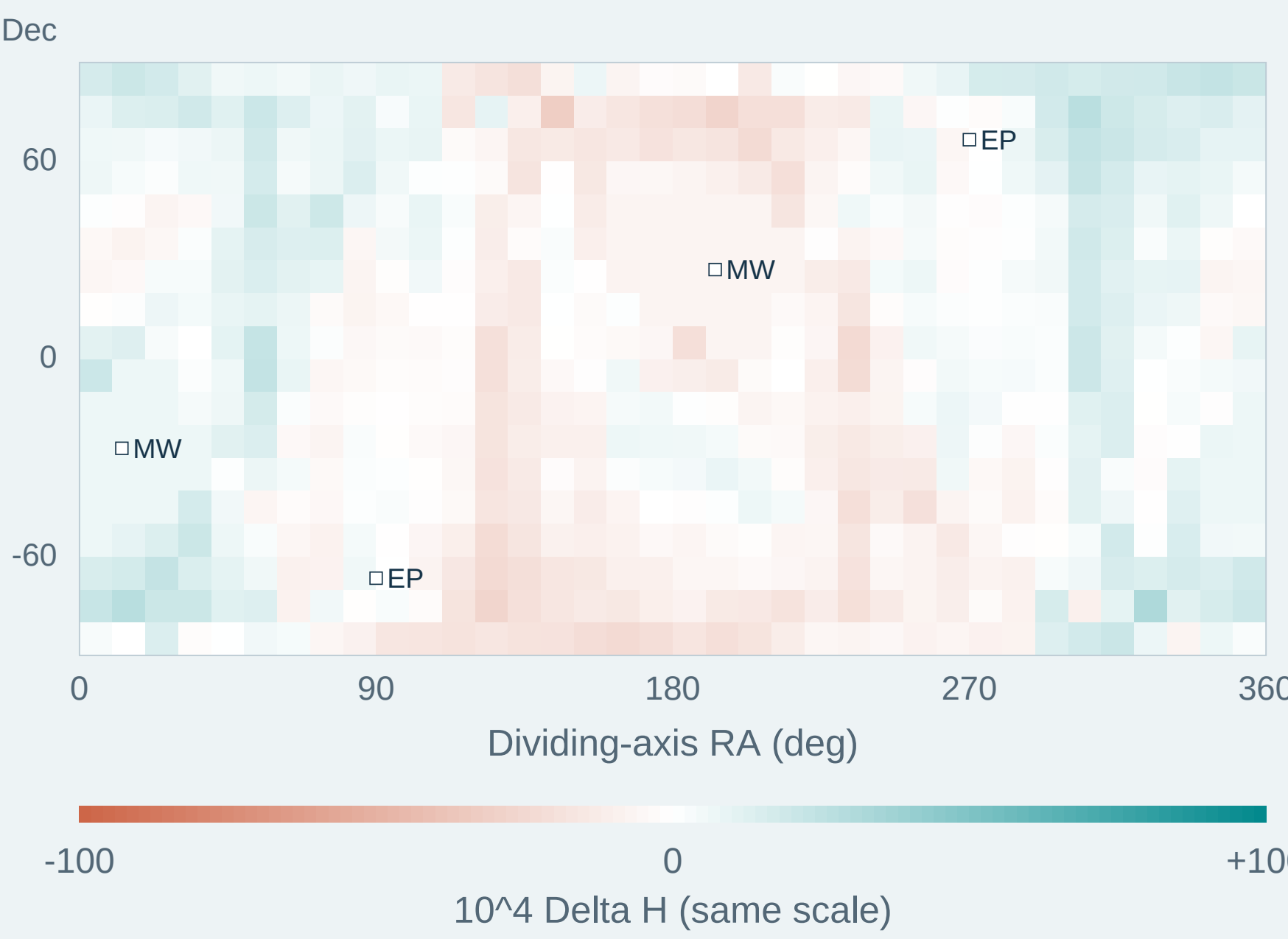
Two halves -> contrast -> direction

H summarizes pair handedness of projected directions. Map colour is $H(\text{toward}) - H(\text{opposite})$.

H is dimensionless, not speed or CW/ACW %.

1 Regression

1,713 groups + 9 clusters after screening



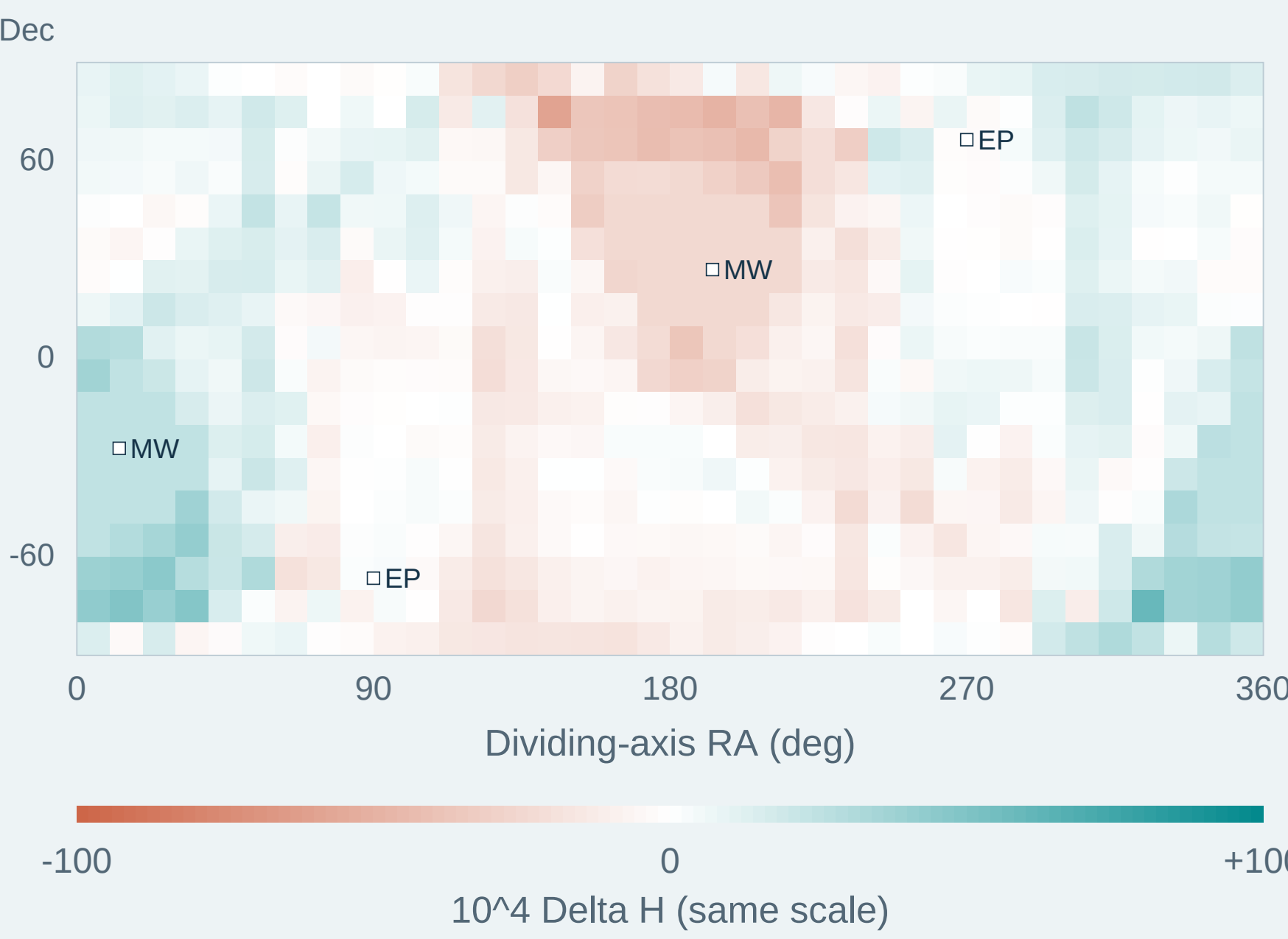
All-system $10^4 H = +0.674$

Leave-one range: +0.023 to +1.319

Sensitivity, not a confidence interval.

2 Half-plane

1,713 groups + 9 clusters after screening



All-system $10^4 H = +0.211$

Leave-one range: -0.416 to +0.839

Sensitivity, not a confidence interval.

Take away

- No statistically significant directional winding pattern: all four new Fourier tests have Bonferroni $p = 1.0$.
- The maps still show how observed imbalance changes with the choice of sky-dividing axis.
- Reference calibration is sample-dependent. Overlapping targets and methods are not independent evidence.
- Groups / clusters: descriptive motion patterns; no new global test or confirmed physical spin.



Research details
NebulaMind

Research details and references: nebulamind.net/studies/spin-parity-2026.html

Duho Kim | KAS 2026 Fall | v61, 10 Oct 2026 | Take away in bullet points; measurements and interpretation unchanged.